

Protostar looks like a star, but its core is not yet hot enough for nuclear fusion

Nebula

Protostar

Main sequence stars fuse hydrogen atoms to helium atoms in their cores.



Protostor

Large Star



Small Star

A Red Giant is formed when a star runs out of hydrogen at its core and starts fusing hydrogen into helium outside the core, releasing energy and expanding the star.

As the Large Red Giant condenses, it heats up even further, burning the last of its hydrogen and causing the star's outer layers to expand further.

Protons and electrons

left after a supernova

to produce a very dense neutron star.

are forced to combine,



Red Supergiant

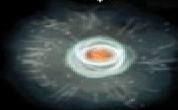


Red Giant

Large Red Giants are hot enough to turn the helium at their core into heavy elements like carbon.



- by the sudden re-ignition of nuclear fusion in a degenera
- by the gravitational collap the core of a massive star.



Once the star runs out of fuel, it collapses under gravity, and the outer lavers are ejected into the vastness of space.

Supernova



Neutron Star



Planetary Nebula



White Dwarf

The remains of the star are devoid of fuel. They consist of degenerate matter with a very high density.

A white Dwarf becomes a black dwarf when it stops emitting light.

If the mass is significantly greater, the gravity will be so strong that the neutron star will shrink further to become a black hole. @ PMF IAS

Birth And Evolution Of Mabive Stars And Stellar Groups

Charles J. Lada, N.D. Kylafis

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Galactic and solar system astrometry Gathered in this volume are the articles they presented Together these provide an auspicious review of optical astrometry and our understanding of the evolution of the Galaxy and solar system for graduate students and researchers The Origin of Stars and Planetary Systems Charles J. Lada, N.D. Kylafis, 2012-12-06 A few years after the publication of The Physics of Star Formation and Early Stellar Evolution we received a request from the publisher for an up dated second edition of this popular reference book As originally intended the volume had proved to be a useful text book for graduate astronomy courses and seminars which dealt with topics related to stellar origins The book was based on a series of lectures delivered by a distinguished group of leading researchers at a NATO Advanced Study Institute ASI held in May 1990 on the island of Crete Greece The primary goal of the ASI was in fact to produce a book which would simultaneously provide a broad and systematic overview of as well as a rigorous introduction to the fun damental physics and astronomy at the heart of modern research in star formation and early stellar evolution However by 1995 concern had arisen among those who used the text as a reference for graduate seminars and courses that the book would need to be updated to stay abreast of the discoveries and progress in this rapidly evolving field After some discussion we concluded that a new edition of the book was warranted and that the goal of producing a new edition would be best accomplished by organizing a second ASI in Crete to review the progress in star formation research Luminous Stars and Associations in Galaxies International Astronomical Union. Symposium, 1986-06-30 Proceedings of the 116th Symposium of the International Astronomical Union held at Porto Heli Greece May 26 31 1985 Neutron Stars: Theory and Observation J.E Ventura, David Pines, 2012-12-06 Some twenty three years after the discovery of pulsars and their identification as rotating neutron stars neutron star physics may be regarded as comingofage Pul sars and accreting neutron stars have now been studied at every wavelength from the initial radio observations through optical X and ray up to the very recent observations in the TeV region while theorists have studied in some detail relevant physical processes both outside and inside neutron stars As a result comparison of theory with observation provides a test of our theoretical ideas in fields as diverse as neutron and nuclear matter superfluidity and superconductivity the acceleration of high energy particles and the generation and maintenance of intense magnetic fields For example through observations of glitches and post glitch behavior of pulsars it has become possible to establish the presence of superfluid neutron mat ter in the inner crust of neutron stars and to determine some of its properties while neutron stars in compact binary systems offer one of the most efficient energy generation mechanisms known It is in fact the interactive interpretation of these diverse pieces of information that can lead to major advances in our understanding of the physics of these exotic objects and justifies the characterization of neutron stars as hadron physics laboratories International Aerospace Abstracts ,1993 The Role of Dust in Dense Regions of Interstellar Matter Thomas Henning, Bringfried Stecklum, 1986-12-31 Proceedings of the Jena Workshop held in Georgenthal GDR March 10 14 1986

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